Mars Exploration Program Science Highlights, Status, Updates

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Mars Exploration Program Legacy



Odyssey Science Highlight

The Water Content of Recurring Slope Lineae on Mars

C. S. Edwards and S. Piqueux, GRL, 2016

- Recurring Slope Lineae (RSL) have been interpreted as present-day, seasonally variable liquid water flows; orbital spectroscopy has not confirmed the presence of liquid H₂O, only hydrated salts.
- THEMIS temperature data and a numerical heat transfer model are used to constrain the amount of water associated with the RSLs
- Surface temperature differences between RSL-bearing and dry RSL-free terrains are consistent with no water associated with RSL and limit the water content of RSL to at most 0.5-3 wt%.
- High thermal inertia regolith signatures expected with crustforming evaporitic salt deposits from cyclical briny water flows are not observed, indicating low water salinity (if any), and/or low enough volumes to prevent their formation.
- The RSL-rich surfaces experience ~100K diurnal temperature oscillations, possible freeze/thaw cycles and/or complete evaporation on timescales that challenge their habitability potential.
- The unique surface temperature measurements provided by THEMIS are consistent with a dry RSL hypothesis, or at least significantly limit the water content of Martian RSL.



THEMIS integrated nighttime images. The white box is the location covering the most active RSL region and the black box is the chosen dry regolith reference area. These colorized ΔT images are created by subtracting the average value from the black box from entire image subset. The temperature variation observed between A & B is primarily due to the azimuthal differences between the two analysis areas and is accurately modeled as a dry regolith surface

MRO Highlight - "Young" Valleys and Lakes in Northern Arabia Terra on Mars

A landscape modified by water and ice is an unambiguous marker of past climate

- This area located just south of the dichotomy boundary on Mars has several fresh shallow valleys (FSVs), some of which flowed into and out of large paleolakes. The extent of these systems was described using CTX data.
- Aided by HiRISE data, this drainage area is dated to be much younger (by hundreds of millions of years) than the ancient channels on early Mars.
- Modeling discovery suggests global periods of warming that allowed ice to melt and water to flow on a cold, wet Mars.



Map of study region showing high concentration of FSVs and associated landforms such as deltas. Color base is topography (see elevation scale). Most FSVs (blue lines) stop at the edges of modelpredicted paleolakes (black) and some flow into and out of paleolakes (e.g., see lake "B").

the northern margin of lake "B" (red box in middle panel). Water continued to flow downhill toward "Heart Lake" (see middle panel)



"Heart Lake" (see middle panel). Wilson, Sharon A., A. D. Howard, J. M. Moore and J. A. Grant (2016), A Cold-Wet Mid-Latitude Environment on Mars during the Hesperian-Amazonian Transition: Evidence from Northern Arabia Valleys and Paleolakes, JGR (#2016JE005052R).



MAVEN Science Highlights





- MAVEN launched Nov 2013; inserted into Mars orbit Sep 2014; primary mission from Nov 2014 – Nov 2015; now in extended mission.
- MAVEN spacecraft and instruments are operating nominally and are providing high-quality science data.
- MAVEN is defining the basic characteristics of the Mars upperatmosphere / ionosphere / magnetosphere system, escape rates at the present epoch, and the processes controlling them.
- Results tell us that loss to space was a major mechanism for the changes in the Mars atmosphere and climate through time.
- First set of results released November 5, 2015:
 - 4 articles in *Science;* 44 articles in *Geophysical Research Letters*.
 - >35 subsequent papers published
 - 2016 Senior Review recommended continuation for an extended mission through FY18 that also endorsed
 - Proposed new measurements modes now ongoing
 - Increased collaborations with other missions
 - Continuing observations are allowing us to understand behavior through a Mars year and with variations in the solar-cycle drivers.

MER Exists Marathon Valley

Mission Status Highlights: MER Opportunity (@ Sol 4510)

- Opportunity left Marathon Valley and is driving east and then south along Endeavour Crater rim
 - 50 nominal (90 sol) missions of exceptional performance on Mars!



http://mars.nasa.gov/mer/gallery/press/opportunity/20150324a.html

Marathon Valley





MEP Updates

- All six missions approved for mission extension
 - Odyssey, Opportunity, Mars Reconnaissance Orbiter, Mars Express/AESPERA-3, Curiosity, Maven
- Calendar
 - Our Red Planet Sept. 20-22, 2016
 - MEPAG virtual Meeting, Oct. 6, 2016
 - AGU public lecture Dec. 11, 2016
 - Landing site Workshop Feb 8-10, 2017
- Mars Program
 - MDAP proposals: Step-1 in Aug 26; Step-2 due Oct. 28, 2016
- Mars 2020 Contamination Control and Planetary Protection Working Group
- Planetary Protection Technology Definition Team

Mars Data Analysis Program

- Currently 108 active grants, and 10 active NASA Earth and Space Science Fellowships, which fund graduate students.
- The August 26 MDAP Step 1 proposals are 166, an increase from previous years (134 and 139 the prior two years).
- Part of the increase this year may have been due to MAVEN, and to the explicit call for HEND proposals.

Contamination Control and Planetary Protection Working Group

Achievement of the stringent levels of organic and biological cleanliness of Mars 2020 returnable samples presents a unique and complex challenge.

- This independent panel of experts is formed to provide the Mars Exploration Program with expert insight into the plans, designs, and operational elements of the sample collection system on Mars 2020.
 - The Working Group will examine contamination control, microbiology, curation, and planetary protection aspects of the sampling system and assess the probable progress toward achieving the level one requirements pertinent to the sampling system and its ability to acquire, cache, and contain pristine samples of Mars.
- The CCPPWG will report to the Mars Exploration Program Lead Scientist and the results relayed to the Planetary Protection Office and the Mars 2020 Project.

Contamination Control and Planetary Protection Working Group

Francis McCubbin – chair JSC Astromaterial curation

Judy Allton

JSC Curation & contamination control

Steve D'Hondt University of Rhode Island Subseafloor microbes

Tony Geller Sandia National Labs Microelectronics contamination

Danny Glavin GSFC Extraterrestrial organics John Priscu Montana State University Polar microbes

Beth Shapiro UC Santa Cruz Ancient DNA

Andrew Steele Carnegie Institution of Washington Extraterrestrial organics

Michael Meyer - convener NASA HQ Mars science/astrobiology/PP

Planetary Protection Technology Definition Team

- Assess technical and engineering challenges to applying available microbialreduction methods, including recontamination prevention, to spacecraft hardware and instruments
- Provide a list of spacecraft and instrument materials known to be compatible with existing planetary protection protocols
- Delineate planetary protection protocols/processes available or which appear promising;
- Identify areas ripe for technological development;
- Evaluate technical and engineering challenges to ensuring that spacecraft hardware and instruments can meet organic cleanliness requirements
- Propose approaches for mitigating the identified challenges, beginning with identification of commonly used materials and spacecraft hardware that are compatible (or particularly vulnerable) to planetary protection protocols
- Identify engineering, technology, and scientific research and development that could be funded by NASA to provide future capabilities to field scientific instruments and spacecraft on missions that require either subsystem or systemlevel microbial reduction and recontamination prevention.

Planetary Protection Technology Definition Team

Members

John Rummel Pat Beauchamp Nancy Carosso Megan Casey Peter Doran Ralph Lorenz **Betsy Pugel David Suarez David Steinfeld** Michael Meyer Quang-Viet Nguyen

<u>Role</u> Chair Technology Contamination Radiation Polar Astrobiology Environment Systems **Rad-Hard ASICS** Thermal Ex Officio Ex Officio

Why More Science?

Mars discoveries leave many mysteries

- Did Mars ever have life? Is it still there?
- Recurring Slope Lineae What are these seasonally changing streaks?
- Methane How much? Does it really come and go? What is the nature of the source (biological or geochemical)? How can it disappear quickly?
- The great transition from a much wetter environment to the cold, dry, acidic planet of today – How, when, and how often did that happen?
- What is the nature of accessible water/ice on Mars? Can it be used?
- Can humans live on Mars? Where are the resources? What are the hazards?