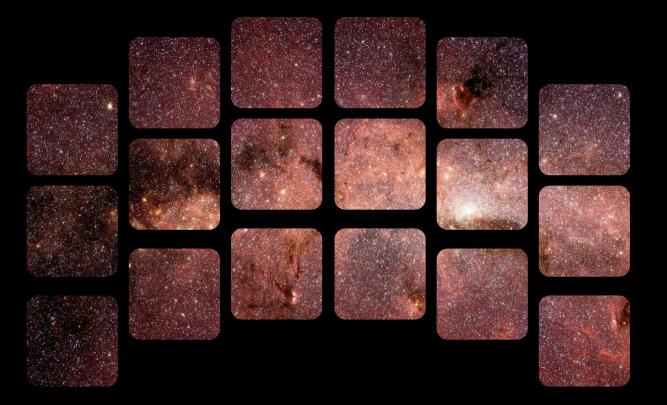


ROMAN⁻

Project Status

Julie McEnery Roman Senior Project Scientist



SPACE TELESCOPE



Overview

- Last presentation was in early March 2020, WFIRST had just entered phase C
- LOTS has happened since then....

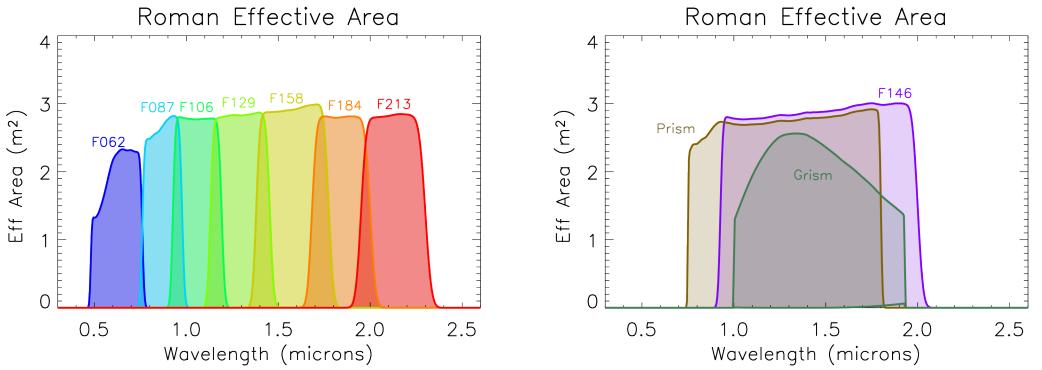


WFIRST is now the Nancy Grace Roman Space Telescope





Added F213 filter (1.95-2.3 micron) – now have imaging filters covering entire spectral range supported by mirrors/detectors!



- 8 imaging filters
- Spectroscopy via prism (0.6-1.8µm, *R*~100, ~24AB) and grism (1.0-1.9µm, *R*~600, ~22AB)



Roman Mission Objectives

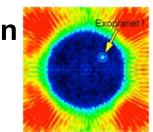
- Wide Field Infrared survey
 - Imaging and spectroscopy to >26.5 AB mag
- Expansion history of the Universe
 - Using supernova, weak lensing and galaxy redshift survey techniques

Growth of Structure in the Universe

Weak lensing, redshift space distortions and galaxy cluster techniques

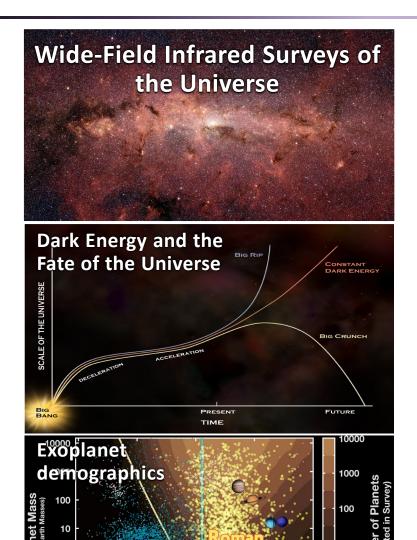
Exoplanet Census

- Statistical census of exoplanets from outer habitable zone to free floating planets
- General Observer program
 - Devote substantial fraction of mission lifetime to peer reviewed general observer program
- Coronagraph technology demonstration
 - Demonstrate exoplanet coronagraphy with active wavefront control



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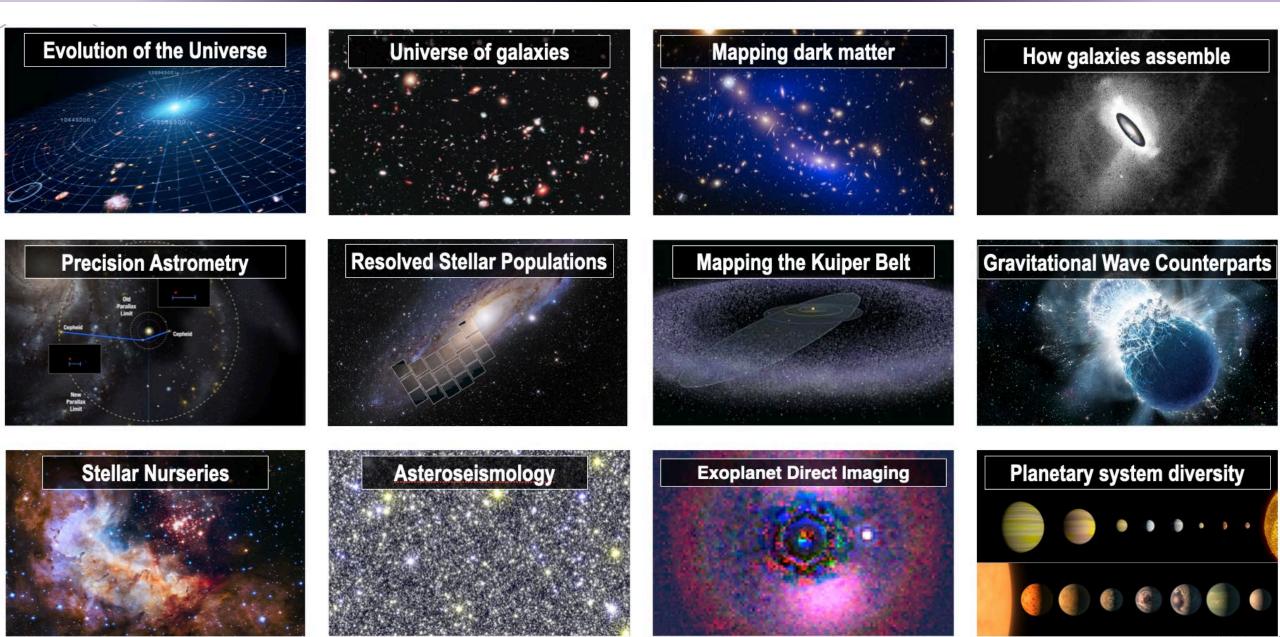
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Semimaior Axis in AU

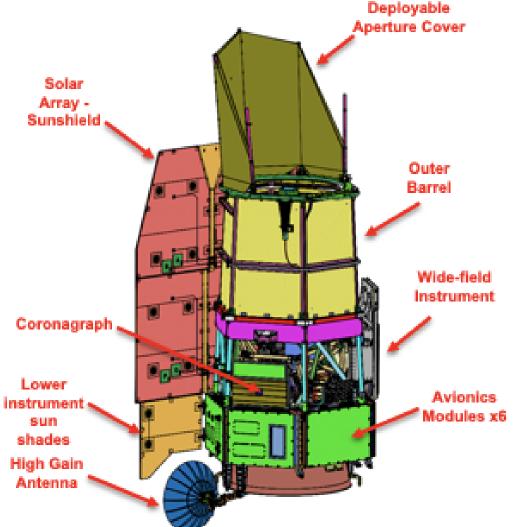
100







Roman Observatory and Instruments



Telescope: 2.4m aperture

Two Instruments:

Wide Field Imager / Slitless Spectrometer

- Vis/Near IR bandpasses (0.48 2.3 micron)
- Field of view 0.281 deg² (~200× HST WFC3-IR)
- 18 4k × 4k detectors (288 Mpixels)

Coronagraph

- Visible bandpass
- Contrast 10⁻⁸-10⁻⁹

Data Volume: 11 Tb/day **Orbit**: Sun-Earth L2

Mission Duration: 5 yr, 10yr goal

https://roman.gsfc.nasa.gov/science/Roman_Reference_Information.html

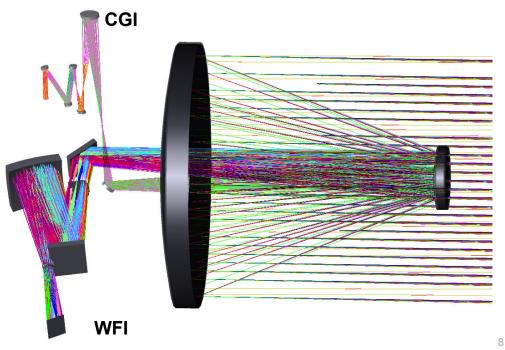


Roman Space Telescope Hardware Status

• Telescope (L3Harris):

- Successful telescope CDR in December 2020
- Primary and secondary mirrors have been refigured, polished and coated
- coronagraph relay optics polished and coated;
- -WFI relay optics
 - Fold mirror 1 final figure complete
 - Fold mirror 2 complete first round of ion figuring
 - Tertiary mirror competed two ion figuring cycles







- Procurement of flight subsystems well underway
- Mechanical Hardware Engineering Development Units (EDUs) nearly complete
- Successful cold deployment test of EDU deployable aperture cover











line to the second seco

Roman 1.7-meter Dual-band High-Gain Antenna (HGA) Engineering Development Unit (EDU) in Goddard's anechoic chamber

Solar Array Sun Shield (SASS) Engineering Development Unit

stowed

Deployed

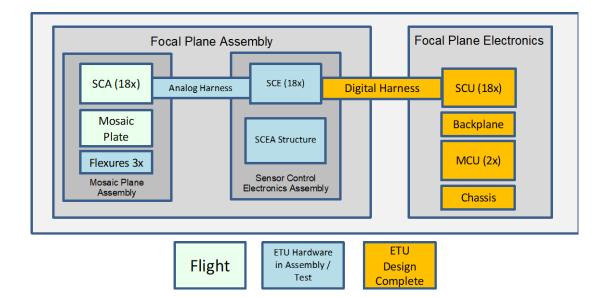


Wide Field Instrument Status

- Wide Field Instrument (GSFC/Ball):
 - Passed WFI CDR!
 - Sensor Chip Assemblies in flight production
 - 20 of 18 flight SCAs accepted
 - 25 more in production, 4 more spares needed

Performance Parameter	SCA Requirements	Current Range of Performance
QE	>80%	86-96%
Mean Dark Current	<0.05 e ⁻ /sec	<0.005 e ⁻ /sec
Total Noise (mean)	<6.5 e⁻ rms	4.8-6.2 e ⁻ rms
Inter-Pixel Capacitance	<2%	1.3-2.0%
Operability	>95%	95.9 - 99.1%
High Intensity Persistence	< .5 e ⁻ /sec	0.06 – 0.21 e ⁻ /s
Low Intensity Persistence	< .15 e ⁻ /sec	0.03 – 0.12 e ⁻ /s

SCA Performance



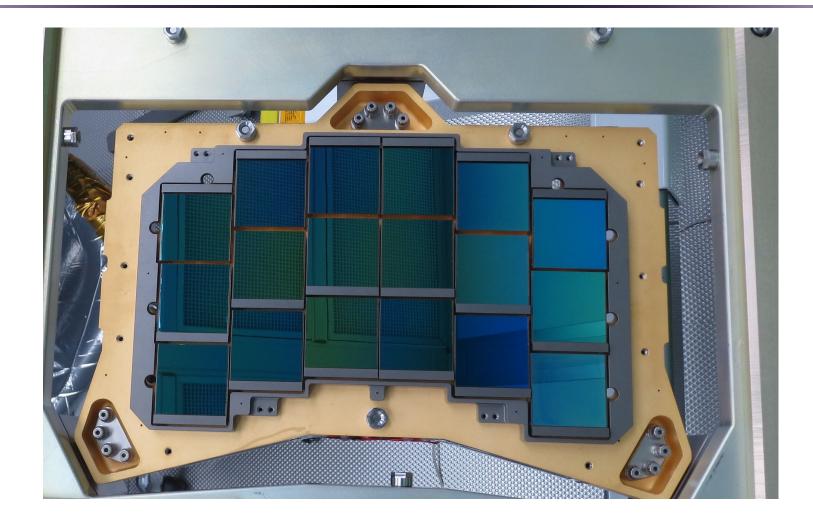
SCE Performance (SCE EDU & ETU Testing)

Performance Parameter	SCE Requirements	Current Range of Performance	SIDECAR Performance
Bias Noise	< 10 µV	< 10 µV	< 102 µV
ADC Noise	< 125 μV	< 82 μV	< 148 µV
Input Noise Rejection	> 13 dB	> 80 dB	-
Power Dissipation	<mark><600 mW*</mark>	<500 mW	

ETU and Flight component and system performance meets requirements. CBE total noise is $6.0 - 6.5 e^{-1}$.



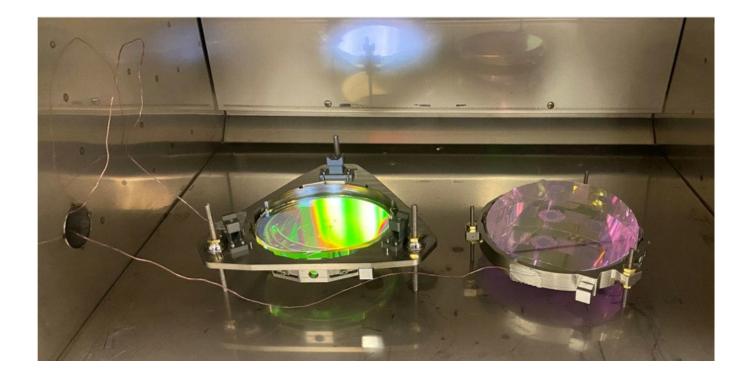
A Pretty Picture!

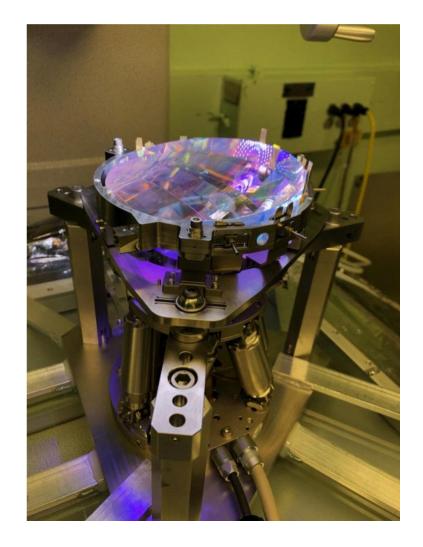


 Completed installation and alignment of all 18 engineering test unit (ETU) sensor chip assemblies (SCA) on the ETU mosaic plate



- Added Kshort science filter
- Flight fabrication underway for grism, prism and filters

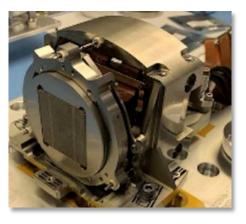






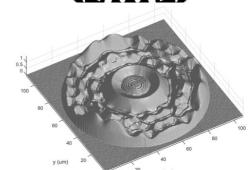
CGI will demonstrate key technologies

Large-format Deformable Mirrors



High-contrast Coronagraph Masks



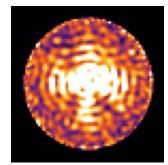


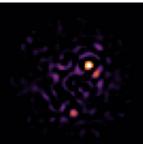
All hardware now at TRL ≥ 6

Ultra-low-noise Photon-counting EMCCDs

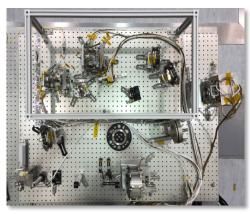


Data Post-Processing





Ultra-Precise Wavefront Sensing & Control (now Ground-In-The-Loop)





Coronagraph Development Status

On track for CGI CDR in April

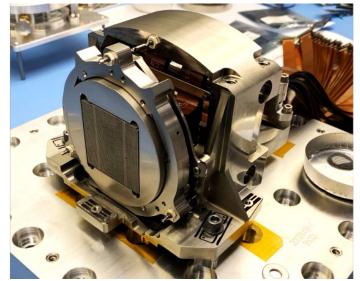
Numerous subsystem reviews complete: DM, II&T, Optical alignment, Mechanical/Thermal, V&V, HOWFS GITL, avionics, camera components

Deformable mirrors

- TRL-6 achieved for deformable mirrors (i.e. passed environmental testing)
- received 2nd flight candidate and spare

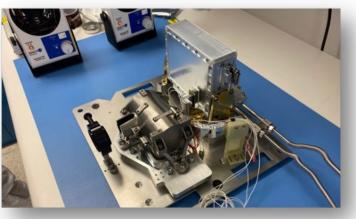
Camera

- Teledyne-e2v (UK) has completed two wafer lots and a third will finish in April. On-track for flightqualified sensor deliveries starting in August.
- ABB/NuVu (Canada) is assembling flight electronics boards and expects to begin testing in April with flight unit delivery planned for August.
- JPL has completed assembly of a complete EDU camera system and is conducting thermal vacuum testing with prototype EMCCD sensors. Flight camera assembly is planned for fall.



TRL-6 DM unit

EDU Camera





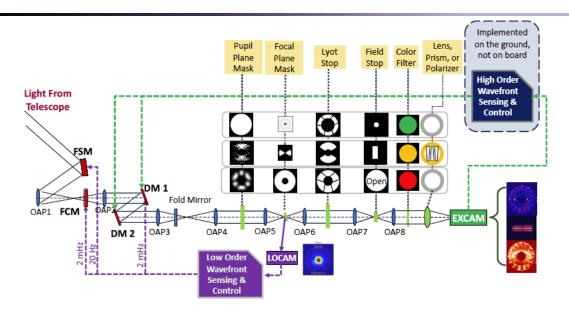
Coronagraph Development Status

Wavefront sensing and control

- Hardware on track
- High order wavefront sensing and control moved from onboard to ground implementation
- Masks
 - Flight masks being fabricated at JPL

Data post processing

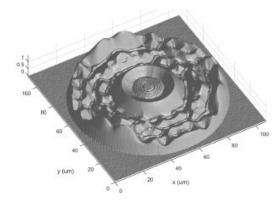
- Simulated imaging and spectroscopy data sets are available on the IPAC Roman website
 - Includes report summarizing processing of the "OS9" simulated imaging dataset
- Tutorials and data challenges available at <u>https://www.exoplanetdatachallenge.com/</u>



Black Silicon reflective shaped pupils

the purple

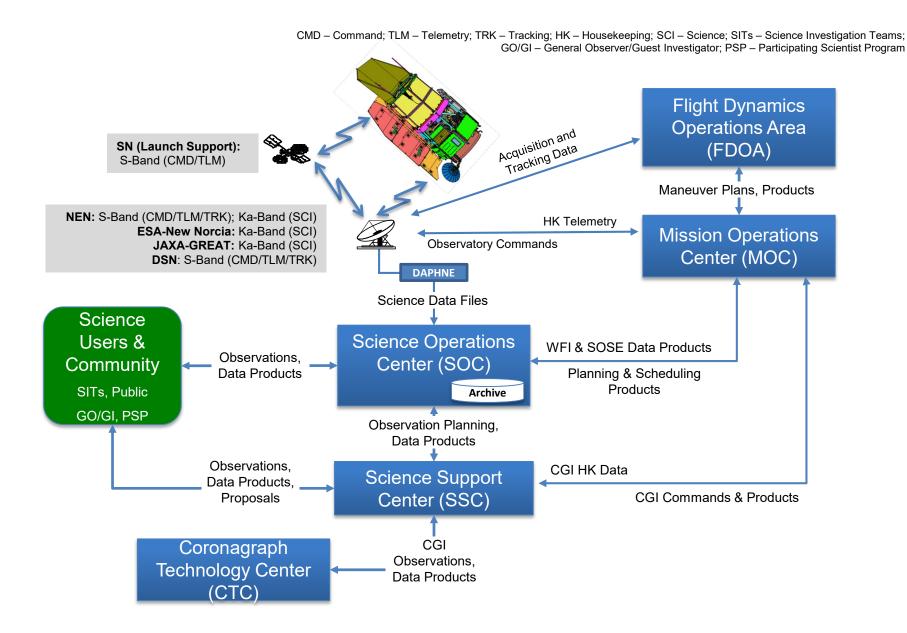
Hybrid Lyot image plane mask



Completed assessment of CGI tech demo value in the perspective of future direct imaging missions (Mennesson et al. 2020,<u>https://arxiv.org/pdf/2008.05624.pdf</u>)

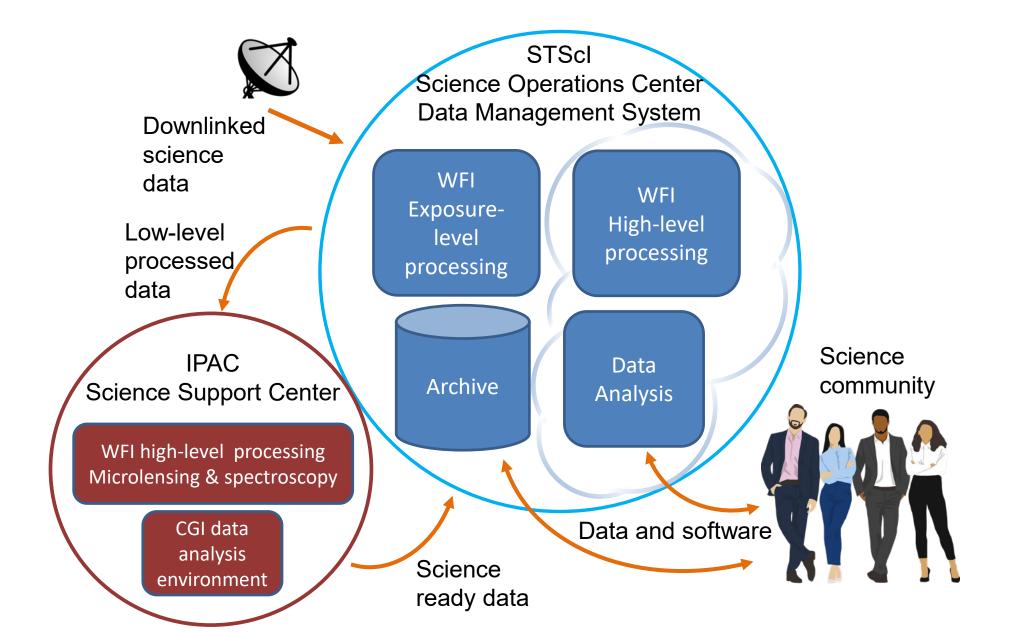


Ground System





Roman Science Operations





- Data system consists of:
 - Pipeline for low level data processing
 - Pipeline for high level processing
 - Science platform (HLPP) allowing users to interact with data and high-level processing software in the cloud
 - Archive with HST/JWST/MAST like functionality
- Updated data processing plans at STScI (SOC) and IPAC (SSC) to augment high level science processing
 - Expanded catalog functions to be implemented by science centers: deblending, photo-z, some time-domain functions etc
 - Astrometry functions e.g. tying to Gaia
 - Some elements of PSF characterization
 - Instrument simulations
- Evaluating science platform and cloud computing options and solutions
 - Including considerations of interoperability with other facilities (e.g. Rubin)



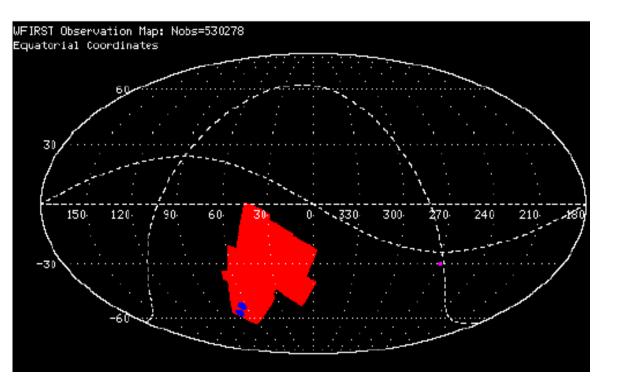
- Demonstrated prototype data platform to science teams last March
 - Opportunity to explore the system for usability, look and feel etc
- Computing and data resource allocation
 - SOC established an internal committee to evaluate computing needs for general Roman users
 - Multi-tiered system operated by the SOC
 - -Access guaranteed for all approved GI/key projects
 - -Simple process to request access by general community
 - -Quotas to control costs within preset bounds
 - Funding via the SOC contract, not GI grants
 - Sized/scoped to enable generic projects by a broad community



- Core Community Surveys: a significant fraction of the prime mission used for revolutionary surveys of unprecedented scale
- Three Core Community Surveys to address 2010 Decadal Survey science goals
 - Extragalactic Wide Area Survey
 - Extragalactic Time Domain Survey
 - Galactic Time Domain Survey
- The definition of core community surveys will be established via an open process, with a goal of maximizing the overall science return while simultaneously meeting the cosmology and exoplanet science requirements
- Several calls for GO surveys closer to, and after, launch
- No proprietary period for any Roman data



Design Reference Mission: Roman Surveys



Galactic Time Domain Survey: 2 deg^2 , 15 min cadence with W filter, 12 hour cadence for R or Z and Y or J for 60 – 72 day seasons; 6 seasons

Extragalactic Time Domain Survey: 16 deg², (wide), 5 deg² (deep), 4 filters (R, Z, Y, J – wide) /(Z, Y, J, H – deep), 5 day cadence, and prism spectroscopy

Extragalactic Wide Area Survey: 2000 deg² (wide), 20 deg² (deep), 4 filters (Y, J, H, F) for wide and deep fields and grism spectroscopy

General Astrophysics Surveys: TBD via proposals

		EWA Wide	EWA Deep	ETD Wide	ETD Deep	GTD
5 sigma J-band sensitivity	Per Visit	-	-	25.5	26.6	24.9
Sensitivity	Integrated	26.95	28.2	28.3	29.4	27.6 (60d)



<u>https://roman.gsfc.nasa.gov/science/rsig.html</u>

- Meeting presentations and notes available on the meetings tab
 - Recent discussions have been on the observing program
 - Currently discussing white paper call to solicit science case for predefining up to one month of observations to be executed within the first two years of the mission
- Annual opportunities to join this group

Megan Donohue (Chair)	Michigan State U.
Zeljko Ivesic	U. Washington
Jessica Lu	UC Berkeley
John MacKenty	STScl
Ashley Villar	Columbia U / Flatiron Institute
Alice Shapley	UCLA
Keith Bechtol	UW, Madison
Saurabh Jha	Rutgers U
Peter Melchior	Princeton U
Dara Norman	NOIRlab
Jessie Christiansen	NEXSci/ CalTech
Rachel Bean	Cornell U
Ryan Hickox	Dartmouth
Lisa Storrie-Lombardi	Las Cumbres Observatory
Dimitri Mawet	CalTech



Science Investigation Teams

- Supernova Cosmology: Ryan Foley, Saul Perlmutter
- Nearby Galaxies: Ben Williams
- Extragalactic: Brant Robertson
- Weak Lensing and Galaxy Redshift Survey: Olivier Dore
- Exoplanet Coronagraphy: Bruce Macintosh, Margaret Turnbull
- Archival Research: Alexander Szalay
- Cosmic Dawn: James Rhoads
- Exoplanet Microlensing: Scott Gaudi
- Milkyway: Jason Tumlinson
- ~300 scientists in total
 - scientific performance requirements related to the specific science area,
 - design of overall observational strategy concept,
 - science data analysis techniques,
 - ground and space calibration requirements,
 - science simulations, precursor observations,
 - ground calibration, observational needs, data processing, ancillary data collection/incorporation, analysis, dissemination and documentation of the proposed science investigation.
- Current science team contracts expire later this year

Adjutant Scientists David Spergel - WFI Jeremy Kasdin - CGI



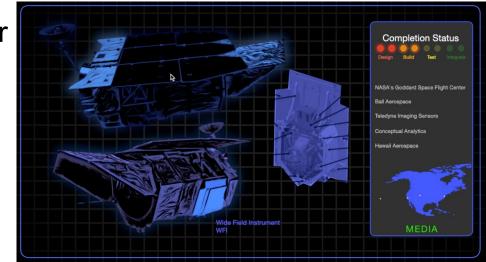
- Roman opportunities announced in the ROSES call last month
 - Currently a "stub" with details to be provided in an amendment
 - Includes opportunities for CGI (CGI Participation program), WFI preparatory science, and key project teams.
 - Accommodates stable long-term funding to support development of needed infrastructure
 - flexible shorter-term opportunities to allow us to be more responsive to a changing science landscape
 - and will allow a variety of different science community models large open consortia, small PI led teams etc



- Coordinated via monthly meetings with all mission partners
- Regular (~monthly) web/press releases
 - "Primary Mirror for NASA's Roman Space Telescope Completed" released on Sept 3
 - "NASA's Roman Space Telescope to Uncover Echoes of the Universe's Creation" released on Nov 18
 - Press release from STScl "Roman Space Telescope Could Image 100 Hubble Ultra Deep Fields at Once" on Jan 11
 - "NASA's Roman Mission Will Probe Galaxy's Core for Hot Jupiters, Brown Dwarfs" released on Jan 25.
- Active social media presence
 - -@nasaroman (twitter, facebook)
- Developing interactive infographic









https://roman.ipac.caltech.edu/Lectures. html

- Monthly lecture series jointly run by IPAC, STScI, JPL and GSFC
- Please join us
- Speaker suggestions welcome

roman.ipac.calte	ch.edu/Lectures.html		☆ ○ \$
Nancy G	ace Roma	n Space Telescope at	IPAC
Iome Science D	1911997777777777777777777777777	TALKS & EVENTS PUBLICATIONS CONTACT	
			and a state of the
Roman V	irtual Lectu	ire Series	
These talks are oper Roman mission. All a	to the entire astronomy co are welcome and encourage	thly virtual lecture series is run jointly by JPL, IPAC, Godo ommunity and cover science, engineering, and technolo ed to attend. Talks are ~20 minutes with ~10 minutes fo from 1-1:30 pm Pacific / 4-4:30 pm Eastern.	gy related to the
	ee: Rob Zellem (JPL), Tiffany Malhotra (Goddard).	/ Meshkat (IPAC), Ryan Russell (STScI), Karoline Gilbert (S	STScl), Julie McEnery
Please contact Rob Z Roman Space Telesc		m@jpl.nasa.gov) if you are interested in giving a talk on	work related to the
To receive monthly l	ecture announcements and	d webinar connection information, please subscribe to t	his mailing list.
Upcoming			
DATE	SPEAKER (AFFILIATION)	TITLE (RECORDING)	ABSTRACT
Past			
DATE	SPEAKER (AFFILIATION)	TITLE (RECORDING)	ABSTRACT
• Dec. 17, 2020	Prabal Saxena (GSFC)	Simulating Roman/CGI Observations of the reflected light exoplanet spectra of the bright, nearby exoplanet ups And d	Abs
• Nov. 19, 2020	Dominic Benford (NASA HQ)	Is Nancy Grace Roman the Most Influential Person You've Never Heard Of?	Abs
• Oct. 15, 2020	Marie Ygouf (JPL)	Post-Processing of Roman Space Telescope CGI Data	Abs
 Sep. 17, 2020 	Harry Ferguson	Sort of Near-Field Cosmology with Semi- Resolved Dwarf Galaxies	Abs



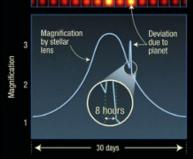
Annual Roman Science Conference

- Hosted in alternate years by STScI or IPAC
- Previous workshops
 - Galaxy Formation and
 Evolution in the Era of the
 Nancy Grace Roman Space
 Telescope (2020),
 - Science in our own backyard – exploring the Galaxy and the local group with WFIRST (2019)
 - Astronomy in the 2020s:
 Synergies with WFIRST (2017)

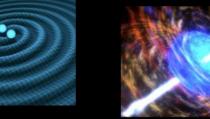
Exploring the Transient Universe with The Roman Space Telescope

..an international meeting to be held in Fall 2021 at IPAC/Caltech









For updates and information see: <u>https://roman.ipac.caltech.edu</u>





- CGI CDR April 2021, Ground System CDR July 2021
- Mission/Spacecraft CDR Sept 2021
- Covid impacts currently rebaselining project
- Roman is on track, and everything where it needs to be approaching CDR
- Exciting to see things coming together