

NASA PCOS X-RAY CONCEPTS STUDY STATUS

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X-ray Concepts Study Scientist



Outline

Background

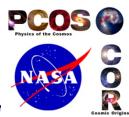
Motivation for Study

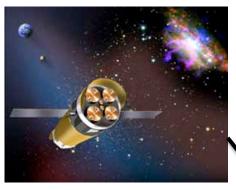
X- ray Study

- Objectives, Process, Participants, Schedule
- Workshop
- NASA Design Lab runs
- Next Steps

Questions/Discussion

The road to the next strategic X-ray observatory



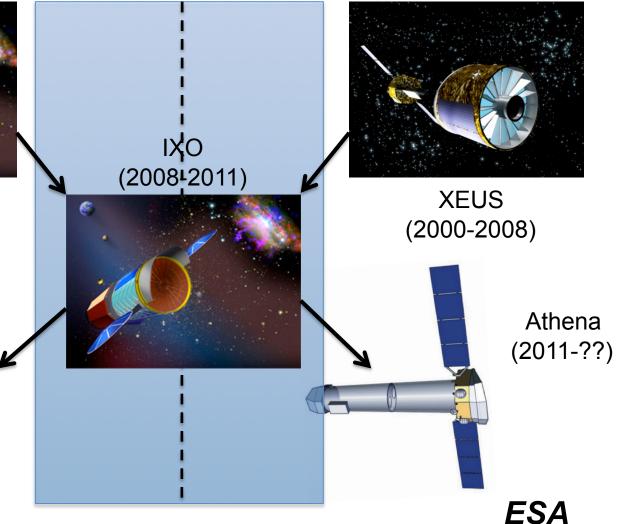


Constellation-X (1996-2008)

- Tech. development plan (2011-2012)
- Mission architecture assessment, with and without US contribution to Athena (2011-2012)
- Mission concept(s) definition (>2012)*

* Purview of the CAA and NASA HQ

NASA





Background behind concepts study

- IXO study activities in US were terminated in fall 2011
 - Prior to termination:
 - Produced mirror development plan consistent with NWNH recommendation
 - Developed AXSIO concept (a descope of IXO)
- X-ray study activities and technology development moved under the Physics of the Cosmos (PCOS) Program Office
 - SAT plus directed support for technology in 2012 and beyond (depending on available budget)
- In September 2011, NASA initiated a concept study to identify more cost effective ways to perform IXO and LISA science



X-ray Concepts Study

Objectives

- Determine the range of science objectives of IXO achievable at lower cost points
- Explore new mission architectures and technical solutions
- Fully engage the community and ensure that all voices are heard, all perspectives considered
- Create data for a report to the CAA that describes options for science return at multiple cost points for X-ray astronomy

Deliver final report to NASA HQ that:

- Describes and analyzes trade space of science return vs. mission cost
- Summarizes the mission concepts developed during the study and how they relate to the trade space and other mission concepts that were not developed in a design lab
- Summarizes the RFI responses and the workshop and describes how they were folded into the whole study



2010 Decadal Science Plan (for X-rays)

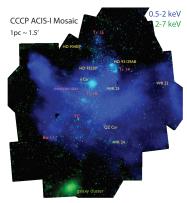
Cosmic Dawn



- Gas exchange between the galaxies and the surrounding inter-galactic medium.
- Formation and growth of black holes in the nuclei of young galaxies

(p.192)

New Worlds

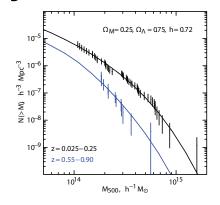


Assess

 habitability by
 using IXO to
 characterize the
 frequency and
 intensity of flares
 on host stars.

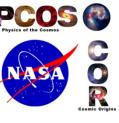
(p.196)

Physics of the Universe



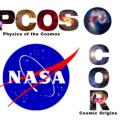
- Find and study distant clusters of galaxies to measure the rate of growth of structure in the universe
- Observe x-rays from gas orbiting close to the event horizon of black holes

(p.200)



Study Phases

- ✓ Request for Information (RFI): ideas for missions and enabling technology. 29 responses received.
- ✓ Community Science Team (CST): 10 members of the community selected by NASA HQ
- ✓ Workshop: provide the community a forum to comment on concepts and technology and identify concepts for further study.
- ✓ Notional Mission Selection: Three mission concepts at different cost points.
- Design Labs: Study team develops concepts through mission design lab runs. Focus is on identifying the technical and cost drivers of each concept.
- Final Report: Summarizes study activities and results for HQ and CAA. Due to NASA HQ on June 7, 2012



Study Boundary Conditions

- The basis for discussion and selection of concepts for further study is the degree of compliance with IXO science objectives, as endorsed by NWNH.
- We are NOT revisiting decadal survey decisions regarding science questions or mission priorities.
- We are studying representative, notional missions for the various cost classes. The goal is to assess the fraction of IXO science that can be performed vs. mission cost. There are no winning or losing concepts.
- No recommendation for a specific mission or a preferred cost class will be given in the final report. This is the CAA's responsibility.
- External constraints (e.g., Athena) will need to be taken into consideration.



Study Team Membership

- Study Manager Gerry Daelemans (GSFC)
- Study Scientist Rob Petre (GSFC)
- Community Science Team 10 members selected by NASA HQ (next slide)
- Science Support Team Andy Ptak (GSFC), Jay Bookbinder, Randall Smith,
 Mike Garcia (SAO)
- Engineering Support Team Tony Nicoletti, Gabe Karpati (GSFC), Mark Freeman, Paul Reid (SAO), discipline engineers
- Support & oversight from:
 - NASA HQ
 - Rita Sambruna (PCOS Program Scientist), Richard Griffiths, Wilt Sanders
 - Jaya Bajpayee (PCOS Program Executive)
 - PCOS Program Office (GSFC)
 - Ann Hornschemeier (Chief Scientist)
 - Jackie Townsend (PCOS/COR Program Office)



Community Science Team Members

- Joel Bregman (Michigan) chair
- Mark Bautz (MIT)
- David Burrows (Penn State)
- Webster Cash (Colorado)
- Christine Jones-Forman (CfA)
- Steve Murray (Johns Hopkins)
- Paul Plucinsky (CfA)
- Brian Ramsey (NASA / MSFC)
- Ron Remillard (MIT)
- Colleen Wilson-Hodge (NASA / MSFC)



RFI responses

- 29 received: 16 mission concepts, 13 enabling technology
 - In the aggregate, the notional missions should probe various points of the science return vs. mission cost trade space.
 - Variety of concepts in nominal "cost bins" (for those with sufficient information for a cost estimate):
 - $1 \le $0.6B \text{ (small)}$
 - 8 ~ \$1.0B (medium)
 - 4 ≥ \$2B (large)
 - Small missions generally addressed one IXO goal well; medium, large mission concepts addressed more than one IXO topic directly
- Technology responses addressed wide range of technology: optics, gratings, calorimeters and other detectors, structures
- All responses posted on PCOS website (http://pcos.gsfc.nasa.gov)



Analysis of RFI responses (concepts only)

Analysis by science support team + CST

- Brief summary of each concept
- Comparison of objectives to those of IXO
- Set of questions created for each concept and sent to respondent
- Responses to questions are being posted

Analysis by engineering support team

- Responses assessed for "completeness" (sufficient information for a Design Lab run?)
- Cost assessed based on information provided (for binning purposes)



Workshop Objectives

- Provide the community an opportunity to comment on the study and shape the missions that will be developed in design labs.
- Provide a forum for discussion and exchanging information between the study team and the community.



Workshop Outcome

- Over 100 participants; more would have attended if timing had been better
- Every respondent to RFI was given an opportunity to present
- General support by community for the current limited study
 - Unclear whether a real opportunity exists, but we need to be ready to take advantage
- Recognition of small number of key technology areas
 - Mirrors (7 concepts used segmented glass)
 - Microcalorimeters (and coolers) (6 concepts)
 - Grating spectrometers (3 concepts)
- Genuine concern about longer term future of X-ray astronomy
 - Competitive strategic mission for 2020 decadal survey
 - Technology funding is marginal at best (APRA, SAT)
 - Is ~\$200M recommended in NWNH going to materialize?
 - Should follow this effort with broader study of future of field (1995 XAPWG)



Notional Missions For Design Lab Runs

- Notion Mission 0: AXSIO
 - Cost goal: < \$2B
 - Formulated and underwent mission design lab run before X-ray study commenced
- Notional Mission I: Gratings only
 - Cost goal: < \$600M
 - Most inexpensive option
 - Concept of choice if Athena is selected
- Notional Mission II: Calorimeter only
 - Cost goal: < \$1B
 - Most desirable scenario if Athena is not selected
 - of three notional concepts, achieves the most IXO science by far
- Notional Mission III: Wide Field only
 - Cost goal: < \$1B
 - Alternative to calorimeter mission if acceptable calorimeter < \$1B not formulated
- All three missions currently configured for an L2 orbit
 - Maximizes observing efficiency (~ 85% vs ~ 65%)
 - Simper spacecraft design
 - Better thermal environment (most relevant for calorimeter mission)



AXSIO: The Advanced X-ray Spectroscopic Imaging Observatory

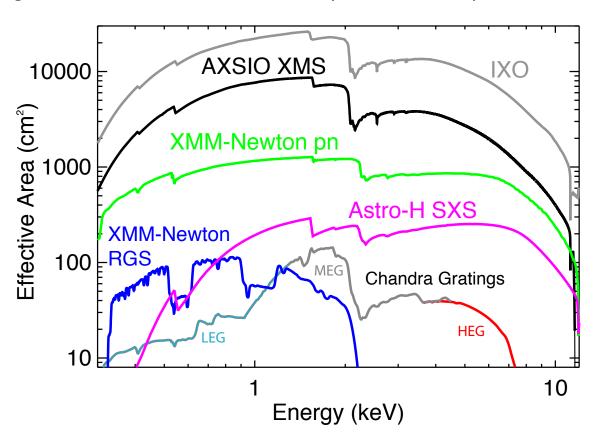
Notional mission "0" designed to meet NWNH recommendations (<\$2B)

Optics: 0.9m² at 1.25 keV; 0.2m² at 6 keV

10" resolution (5" goal)

Calorimeter 40X40 array with < 3 eV resolution

Grating $\lambda/\Delta\lambda > 3000$; ~1000 cm² (0.3-1.0 keV)







Notional Mission I: Gratings

- Merging of off-plane (WHIMEx) and critical-angle transmission (Aegis) designs
 - Takes advantage of segmented mirror design by only using part of aperture (azimuthally) to reduce mass
 - 7 m focal length
- >500 cm² over 0.3-1.3 keV
- R > 3000
- Key science
 - Large-Scale Structure: Warm-Hot Intergalactic Medium
 - Matter under Extreme Conditions
 - Active Galactic Nuclei outflows, absorbers
 - X-ray binary accretion physics
 - Life Cycles of Matter and Energy: Stars and Protostars

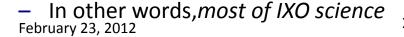


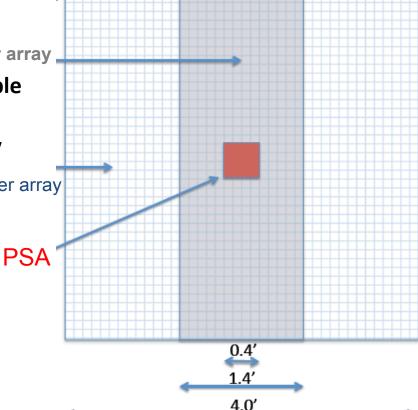
Notional Mission II: Calorimeter

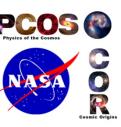
- Descope of AXSIO mirror
 - smaller size and mass
 - 9.5 m focal length

Single pixel outer array

- Central "Point Source Array" (red square) to enable high-count rate science (X-ray binaries)
- Spectral resolution of < 3 eV (inner arrays), < 6 eV (outer array)
- Key science:
 - Matter under Extreme Conditions
 - AGN accretion, outflows
 - X-ray binaries
 - Large-Scale Structure
 - Clusters of galaxies
 - WHIM
 - Life Cycles of Matter
 - Starburst galaxies, Supernova remnants, Stars







Notional Mission III: Wide-field Imager

Focal-plane with CCD-like detectors

- > 24' FOV with ~ 7" spatial resolution constant
- CCD resolution removes high spectral resolution science goals

Two possible mirror designs

- 3 identical mirrors (with three focal planes)
- Single AXSIO-like mirror

Rationale

Alternative if calorimeter mission cannot be brought under \$1B
 without giving up too much essential science and can itself be under \$1B

Key science (done statistically)

- Black hole evolution, high-redshift obscured AGN
- Large-scale structure, clusters of galaxies

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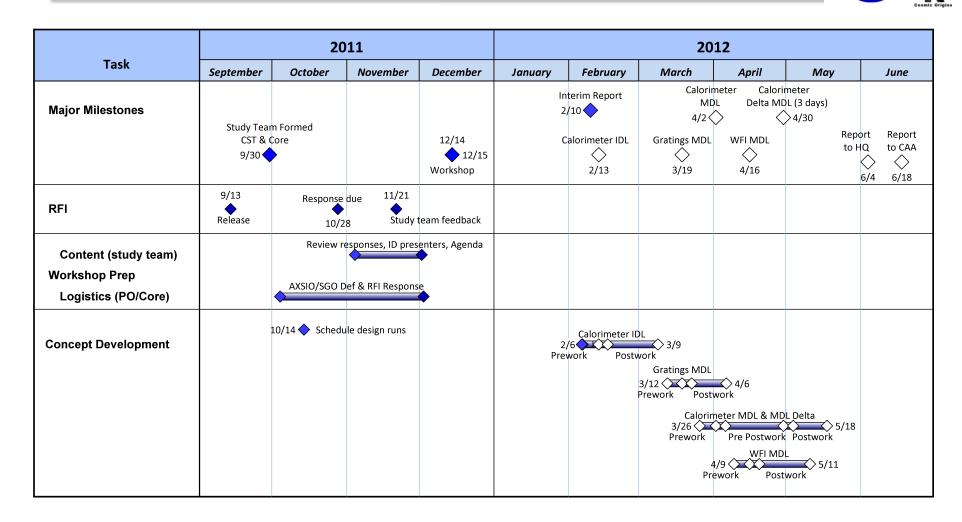


Design Lab Runs

- One Instrument Design Lab (IDL) run for calorimeter
 - Completed Feb. 12-17, final report in ~ 3 weeks
 - Resulted in significant reduction in mass and simplification of focal plane
- Three Traditional Mission Design Lab (MDL) runs
 - Grating mission 3/19-23
 - Calorimeter mission 4/2-6
 - WFI mission 4/16-20
- "Delta"s to each notional mission assessed by MDL personnel
 - Date: 4/30-5/6
 - Plan is to try alternatives to reduce cost
 - Calorimeter: major reduction in mirror, simpler cooling system
 - Grating: LEO instead of L2
 - WFI: Reduce number of telescopes or smaller single telescope design
 - Not all of these deltas can be tried within current time frame
 - May also need MDL time to assess cost to add instruments to singleinstrument designs



Schedule





Enabling Technology

- RFI solicited responses regarding enabling technology
- Technology responses will be used to:
 - Inform discussion about notional missions
 - Provide input to NASA about key areas to be addressed through APRA and PCOS funding
 - Identify in study report critical and enhancing technology areas where support needed for short and long term needs



Next Steps

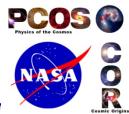
Interim Study Report (done)

- Summarize RFI responses and analysis
- Summarize Workshop Outcome
- Summarize concepts selected for further study
- Design labs (ongoing)
 - Study team is collecting data needed for design lab and defining trades
- Results from design runs plus external trades incorporated into study report
 - Outline completed, CST and study team members assigned writing and editing duties



Involvement of the broader community

- Workshop
- Mostly through the PCOS web site
 - RFI responses and study team summaries
 - Workshop presentations
 - Regular status reports
 - Community "bulletin board" for comments
 - Study report
- Presentations to PhysPAG (Austin and beyond)
- Informal "town hall" at Austin AAS meeting (Tuesday evening)
- Regular progress reports distributed through PCOS, HEAD newsletters
- Final study report will be distributed to the community online and summarized at conferences.



Backup Slides



Near Term X-ray Landscape – a "Golden Age"

Current missions

- Chandra (1999-) Unprecedented imaging and (limited) high resolution spectroscopy
- XMM-Newton (1999-) Wide field, good angular resolution, large area
- Suzaku (2005-) Very broad-band, low surface brightness spectroscopy
- Swift (2004-) Rapid response to transients, deep 10+ keV survey
- RXTE (1996-2012)

Upcoming missions

- NuSTAR (2012) Hard X-ray imaging (6-80 keV)
- Astro-H (2014) Spatially resolved high res spectroscopy + broad band
- GEMS (2014) Sensitive X-ray polarimetry
- Astrosat (2012?) RXTE-like X-ray timing
- SRG (eROSITA, ART) (2013) 0.3-15 keV all sky survey
- NICER (NASA), LOFT (ESA) Timing missions in study phase
- No guaranteed mission after 2017 "golden age" will end abruptly
- What's missing is a strategic mission



Background for concepts study (1 of 3)

- Constellation-X was NASA's strategic X-ray mission concept (1996-2008)
 - High resolution spectroscopy with large effective area
 - Four mirrors (~3m² @ 1 keV) with calorimeter and gratings, plus HXT
 - Ranked 2nd among large missions in 2000 decadal survey
 - Substantial technology development mirrors, calorimeters, gratings
- X-ray Evolving Universe Spectrometer mission (XEUS) was European counterpart (2000-2008)
 - Very large collecting area with wide field imaging plus high resolution spectroscopy, high time resolution and polarimetry
 - One of three L-class mission candidates selected in 2006 for ESA's
 Cosmic Visions
- In 2008, the missions were merged to form the International X-ray Observatory (IXO)
 - Submitted to both US Decadal Survey and ESA Cosmic Visions



Background behind concepts study

- In 2010 IXO was among the four large missions identified in New Worlds New Horizons to move forward this decade
 - Science was very well received but projected cost thought to be too high (\$3.3B-\$4.8B)
 - Key IXO science is high resolution spectroscopy
 - IXO should cost no more than \$2B; 10 arcsec resolution acceptable
 - Strong recommendation to develop optics and other enabling technology to higher readiness level (~\$200M over decade)
- In March 2011, ESA decided to redefine all three L candidates
 - Due to JWST costs and budget realities, ESA was no longer confident that NASA could contribute its share, about half the cost of each mission
 - ESA component of mission to be less than ~€800M
 - US participation possible at ≤\$150M level
 - Downselect decision postponed until February 2012 (now April at earliest)
- ESA decision meant the termination of the formal IXO project (and LISA) by NASA
- ESA has subsequently developed Athena concept
 - Two mirrors, focusing onto WFI and microcalorimeter
 February 23, 2012



The Decadal View of IXO

 "IXO is a versatile, large-area, high-spectralresolution X-ray telescope that will make great advances on broad fronts ranging from the characterization of black holes to elucidation of cosmology and the life cycles of matter and energy in the cosmos. Central to many of the science questions identified by this survey, IXO will revolutionize high-energy astrophysics with more than an order-of-magnitude improvement in capabilities." (p19, emphasis. added)

Broad Range of Mission Concepts submitted as RFI responses



- The Warm-Hot Intergalactic Medium Explorer
- The Advanced X-ray Timing Array
- The High Energy X-ray Probe
- The Black Hole Evolution and Space Time Observatory
- SuperMon & Black Hole Tracker
- An Astrophysics Experiment for Grating and Imaging Spectroscopy
- The Extreme Physics Explorer
- Spectral Analysis with High Angular Resolution Astronomy
- Orbiting Wide Field X-ray Imaging Spectrometer
- Wide Field X-ray Telescope Mission
- Epoch of Reionization Energetic X-ray Survey
- Xenia: A Probe of Cosmic Chemical Evolution
- The Advanced X-ray Spectroscopic Imaging Observatory
- Square-Meter Arcsecond Resolution X-ray Telescope



Calorimeter Mission Effective Areas

